

# Systematics in Intermediate Mass Fragment Produced in Heavy-Ion Collisions\*

Yogesh K. Vermani, Rajeev K. Puri

Department of Physics, Panjab University, Chandigarh-160014, India.

Contact e-mail: [rkpuri@pu.ac.in](mailto:rkpuri@pu.ac.in)

The last few decades have seen advent in heavy-ion research due to production of nuclear matter away from the saturation density. The low density phenomena such as multifragmentation, has drawn great attention as a signature of liquid-gas phase transition. We aim to present here fragment formation as a function of excitation energy for different system masses.

The calculations were made within the framework of *quantum molecular dynamics* (QMD) Model [1]. We simulate the central HI collisions for three entrance channels  $^{129}\text{Xe} + ^{131}\text{Sn}$  ( $E = 45$  to  $140$  MeV/A),  $^{139}\text{La} + ^{139}\text{La}$  ( $E = 50$  to  $120$  MeV/A) and  $^{197}\text{Au} + ^{197}\text{Au}$  ( $E = 70$  to  $130$  MeV/A). Model calculations are done with soft equation of state.

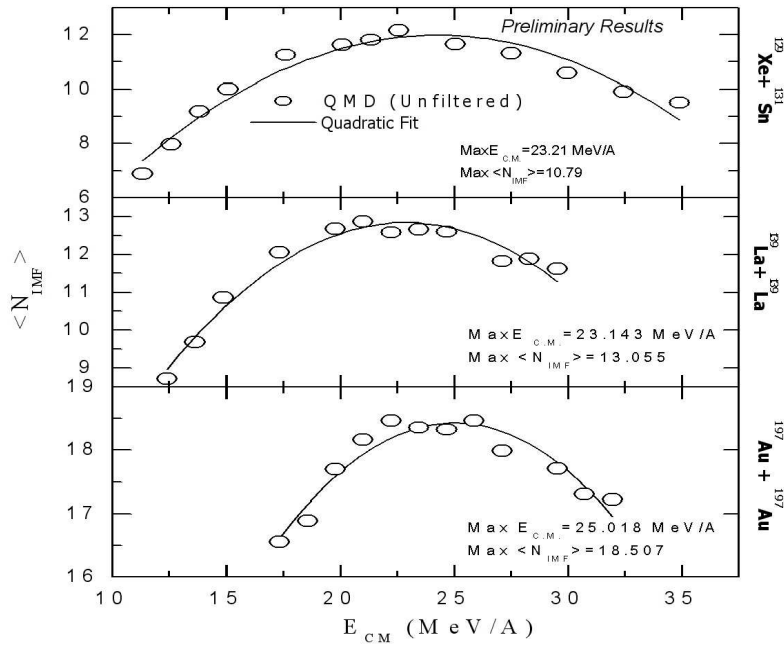


Figure 1: Multiplicity of IMF's as a function of available center of mass energy

In fig. 1 we display the dependence of mean IMF multiplicity  $\langle N_{\text{IMF}} \rangle$  on center-of-mass energy  $E_{\text{CM}}$  (open circles). Solid curves are the quadratic fits to calculated  $\langle N_{\text{IMF}} \rangle$ . Quadratic curves give an estimate of the peak  $\langle N_{\text{IMF}} \rangle$  and the beam energy at which peak occurs. It is clear that (i) systematics of rise and fall in  $\langle N_{\text{IMF}} \rangle$  in these systems is consistent with onset of multifragmentation followed by vaporization at higher beam energies [2], and (ii) the center-of-mass energy available for peak IMF emission shows a weak rising trend with system size. This analysis reinforces the experimental attempts to study fragmentation over wide range of system size.

\* This work is supported by CSIR.

[1] J. Aichelin, Phys. Rep. **202**, 233 (1991)

[2] T. Li et al., Phys. Rev. Lett. **13**, 1924 (1993); G. F. Peaslee et al., Phys. Rev. C **49**, R2271 (1994)