## THE COULOMB DISSOCIATION OF <sup>8</sup>B; A NEW TOOL FOR NUCLEAR ASTROPHYSICS \*

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The Coulomb Dissociation (CD) method was developed over the last twenty years and reached the level of confidence where the very method and its application to the CD of  $^8B$  can be tested with high accuracy. We observe that both the GSI1 [1] and GSI2 [2] measurements of the CD of  $^8B$  are in excellent agreement with the most recent Direct Capture (DC)  $^7Be(p,\gamma)^8B$  reaction measurement performed at Weizmann [3] and in agreement with the Seattle result [4]. We also show [5, 6] that the statements on "fundamental differences" between CD and DC data arise from a misunderstanding (as well as misrepresentation) of CD experiments.

However, in spite of the general good agreement between DC and CD data the slope of the astrophysical cross section factor  $[S_{17}(E)]$  can not be extracted with high accuracy due to discrepancies of the slope measured by Weizmann [3] and Seattle [4] as well as a discrepancy between the slope measured by GSI1 [1] and GSI2 [2]. This discrepancies of the measured slopes (among DC data themselves and among CD data themselves) lead to an additional uncertainty of the extrapolated zero energy cross section factor  $[S_{17}(0)]$  and must be alleviated by future experiments in order to achieve a high precision determination of the extrapolated  $S_{17}(0)$  and the predicted  $^8B$  solar neutrino flux.

- \* Work Supported by USDOE grant No DE-FG02-94ER40870.
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